

Мир двойных звёзд

Константин Маланчев, ГАИШ МГУ

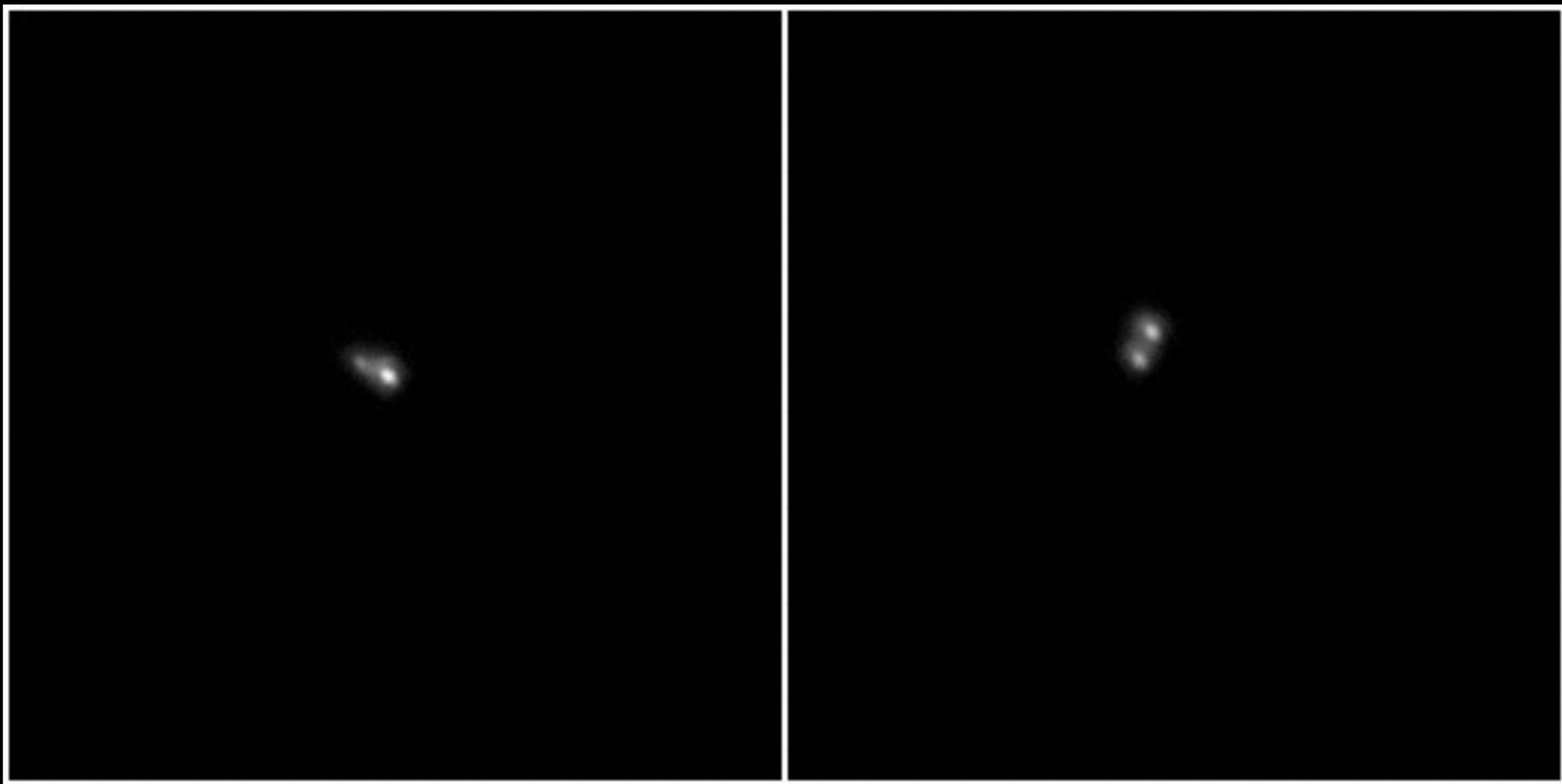


Фото: Сергей Лисаков



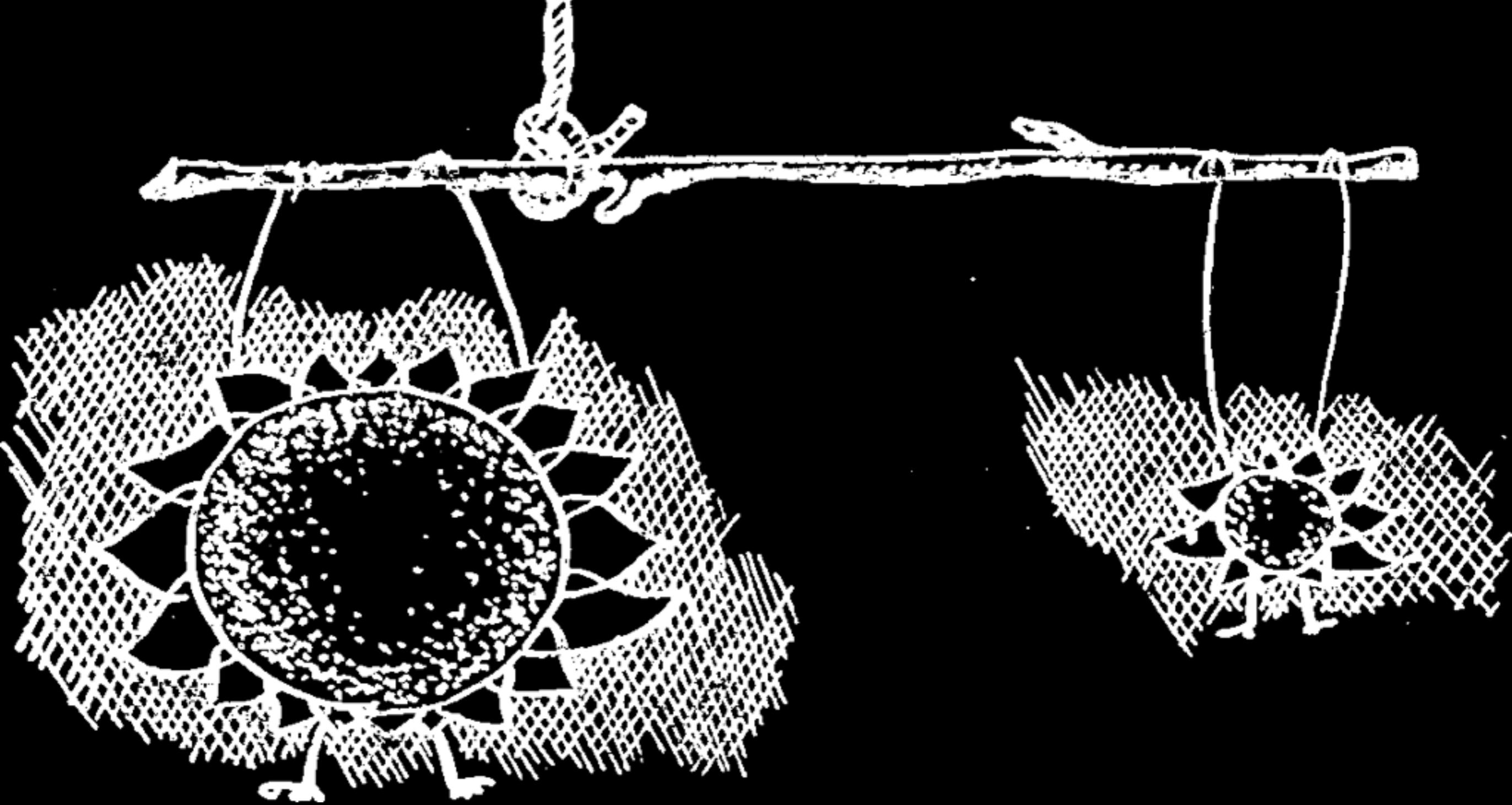
Эпсилон Лирьы

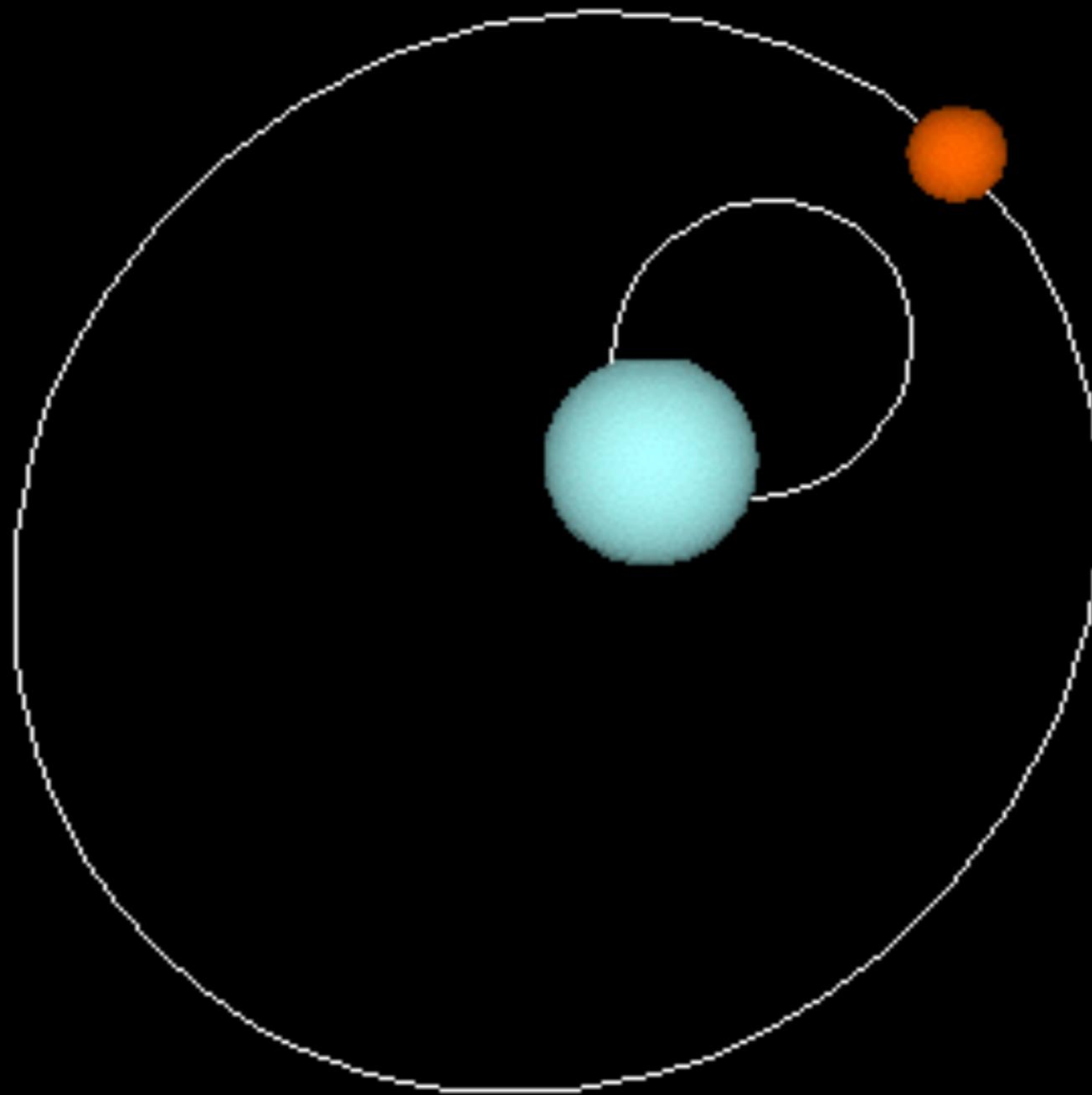
Фото: Сергей Голышев



Эпсилон Лирьы

Фото: Сергей Голышев

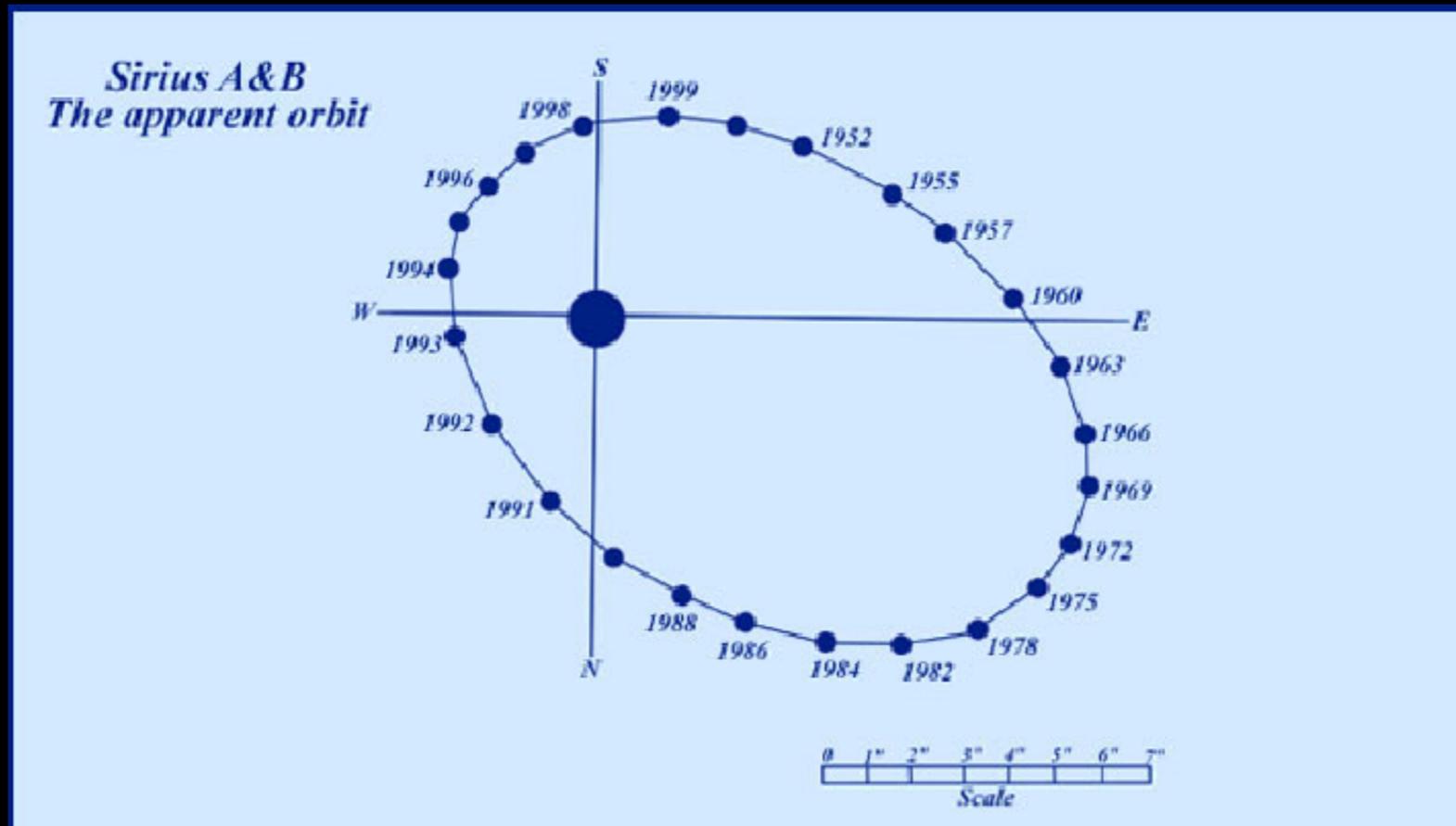
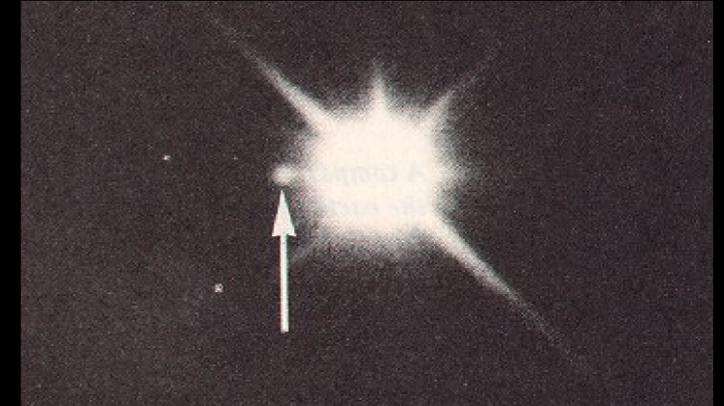
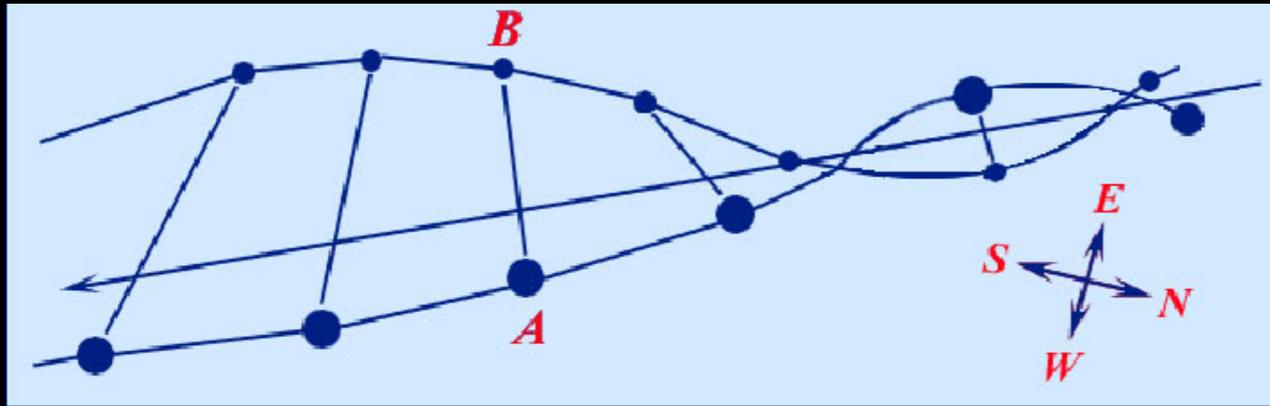




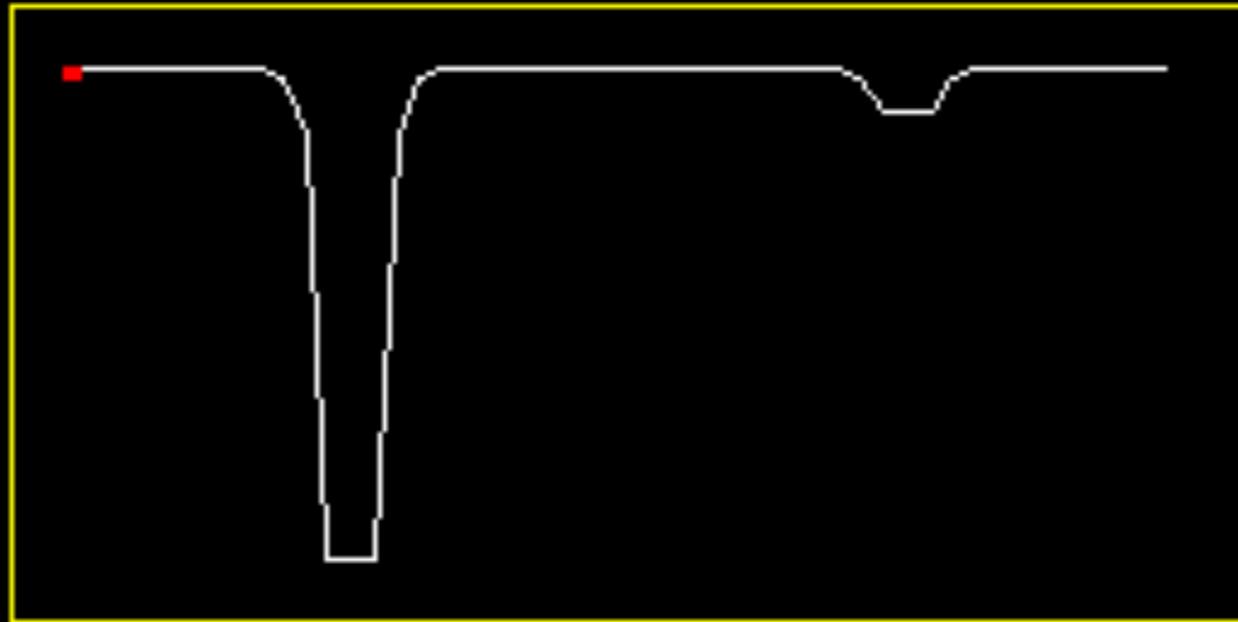
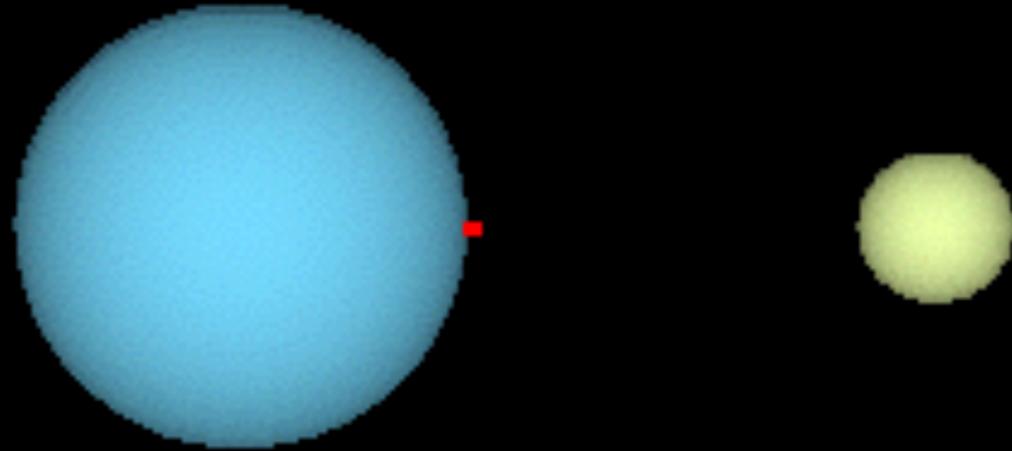
$M_1/M_2=3.6; e=0.4$

Визуально-двойная звезда

Анимация: P. Rogge

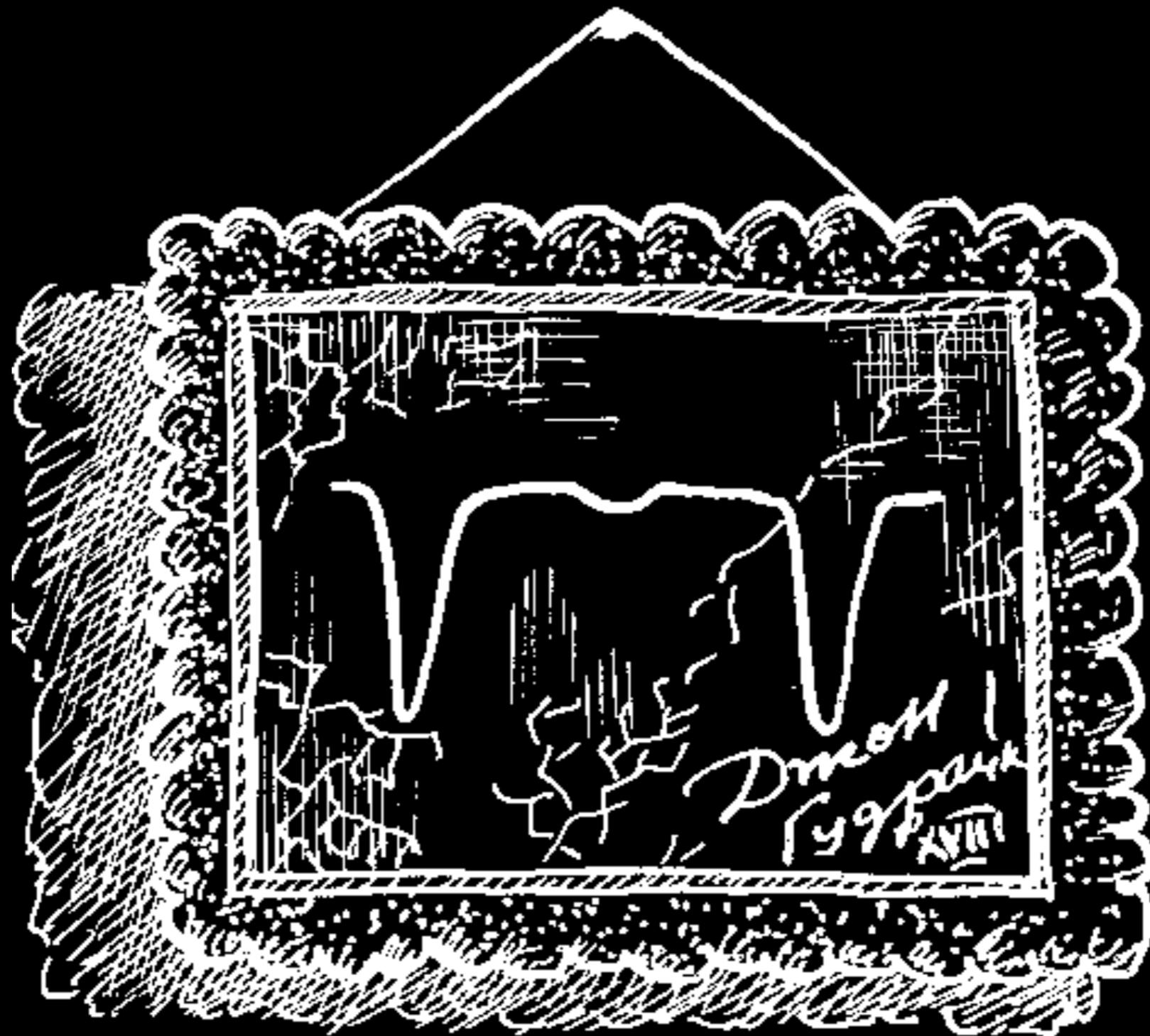


Сириус Б



Затменная двойная звезда

Анимация: P. Rogge



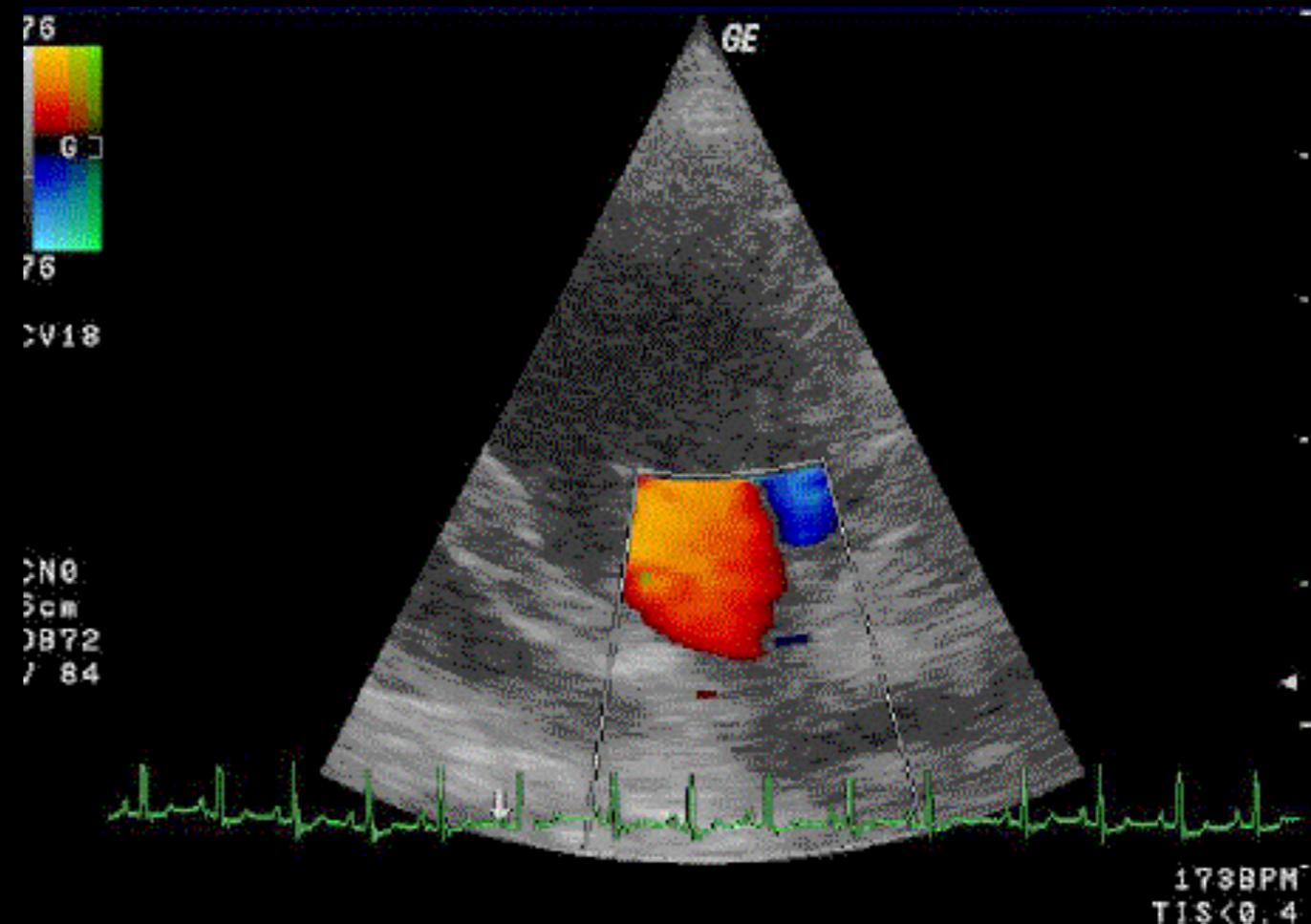
β Ρετζ

АЛГОЛЬ

Изображение: В. Липунов



Фото: Gian Wallace

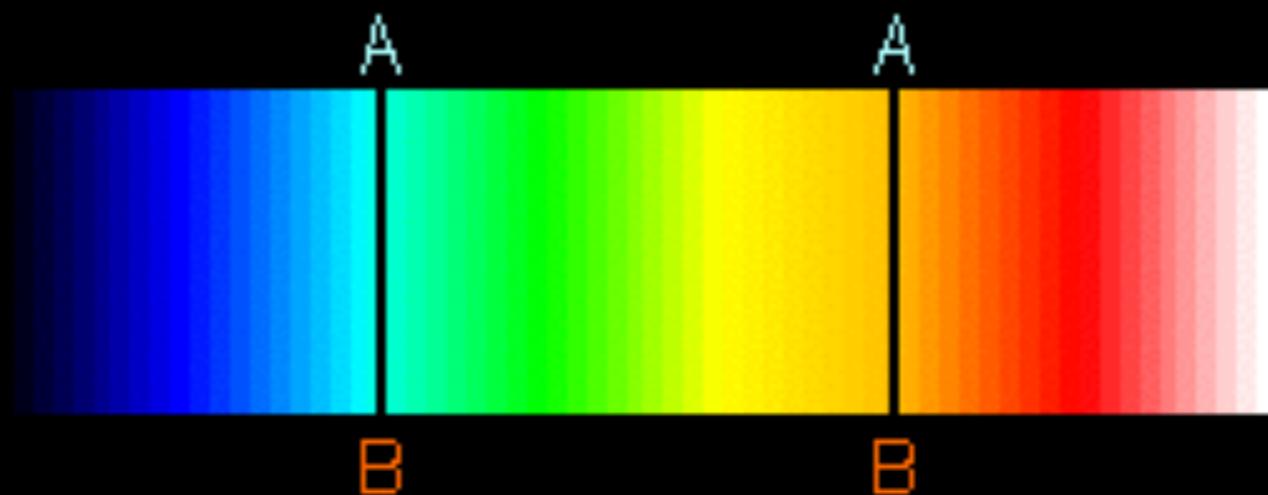
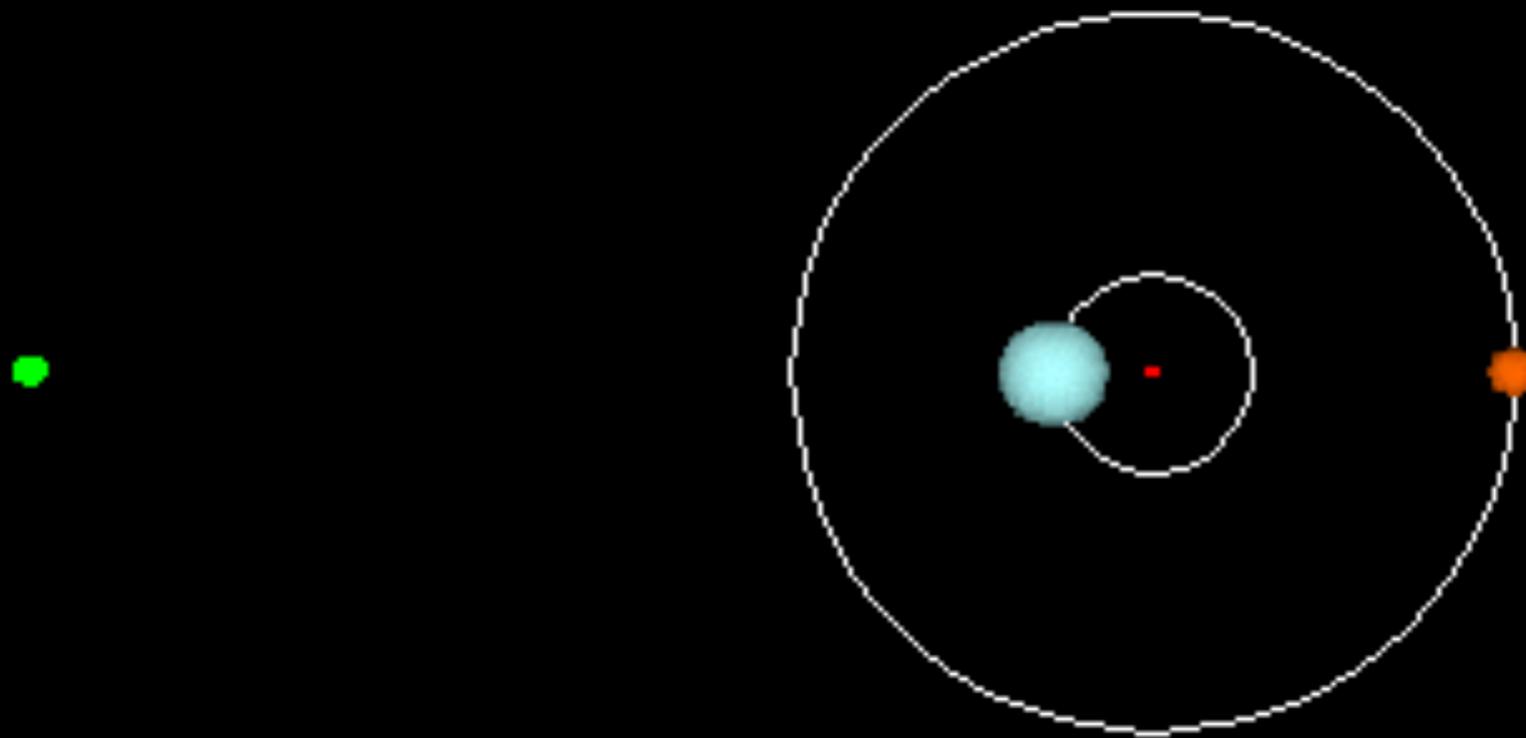


Анимация: Kolumet, Wikipedia Commons



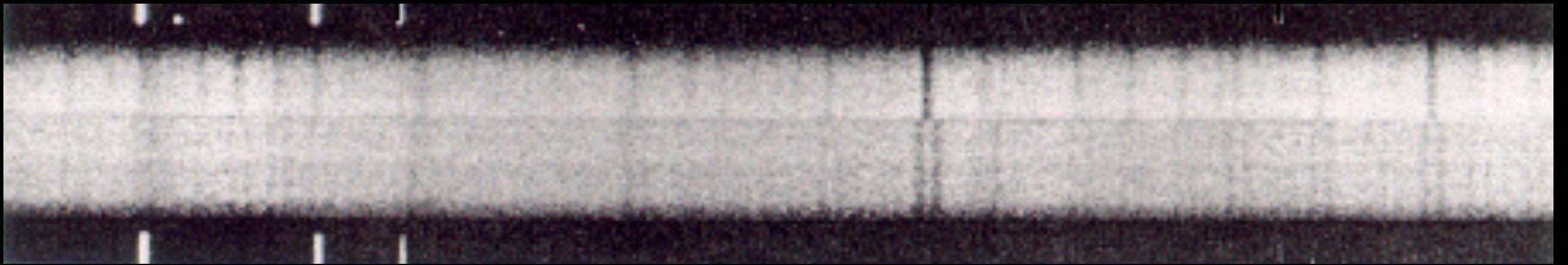
Изображение: TxAlien, Wikipedia Commons

Эффект Доплера

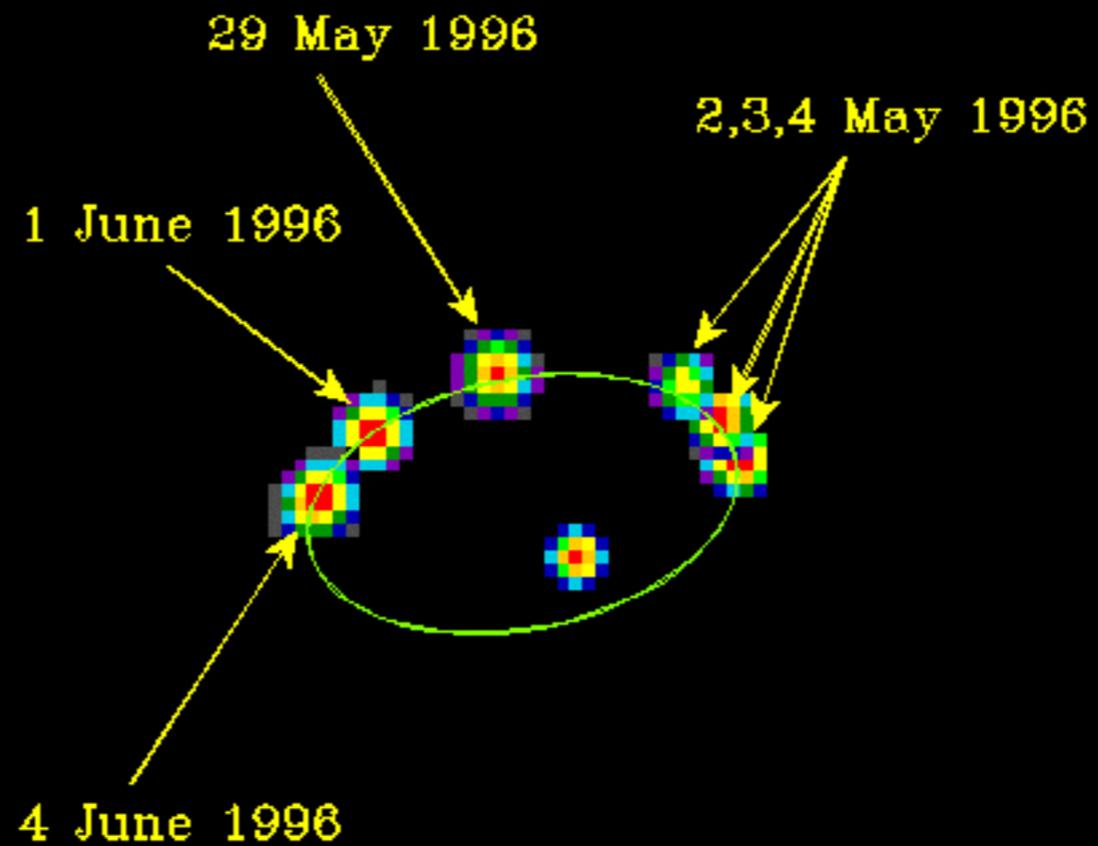


Спектрально-двойная звезда

Анимация: P. Rogge



Изображение: Caltech



Изображение: R. Nemiroff, J. Bonnel

Мицар А

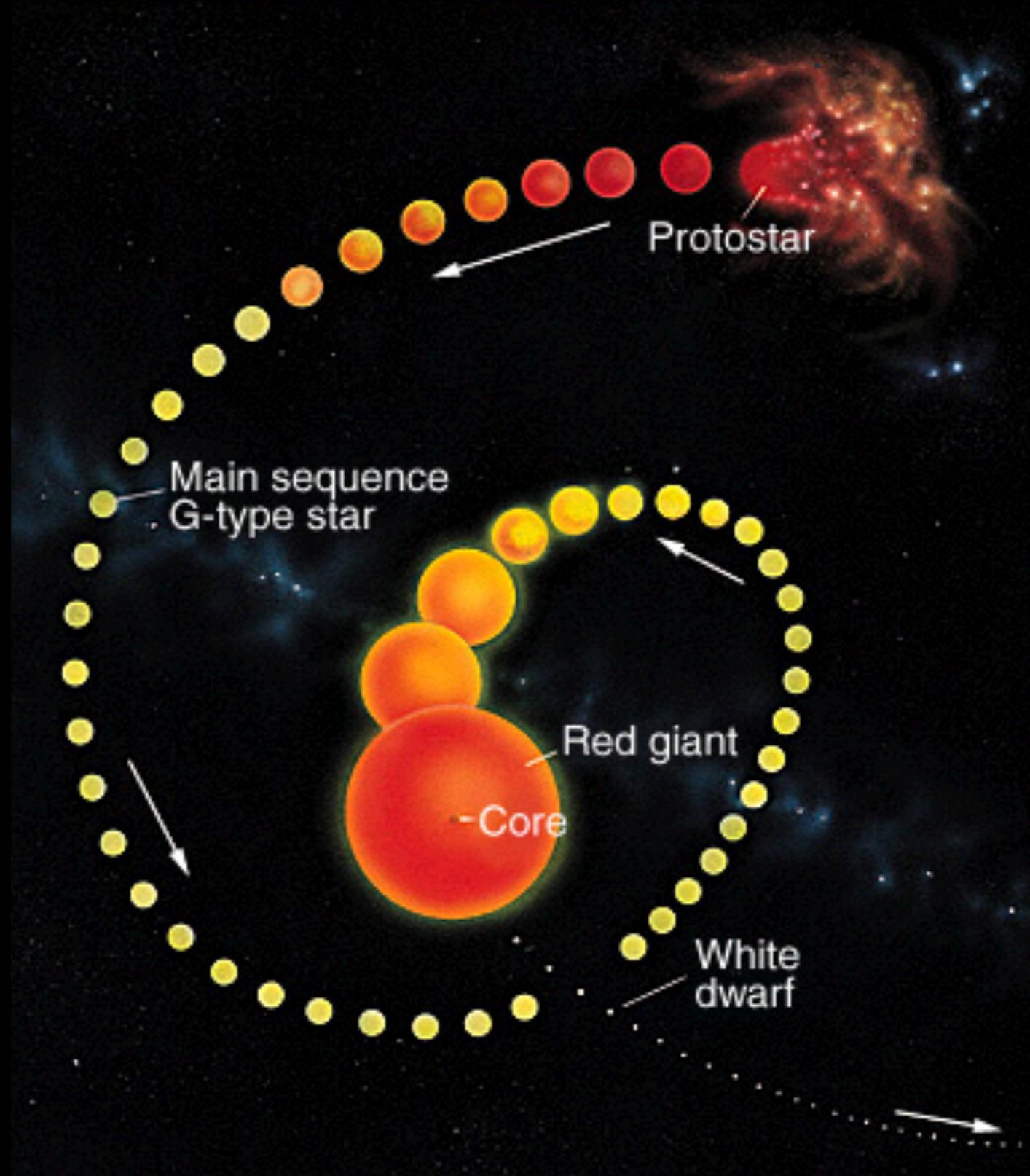
$Z=0.1 Z_{\odot}$

228212 yr

<http://www.ukaff.ac.uk/starcluster/>

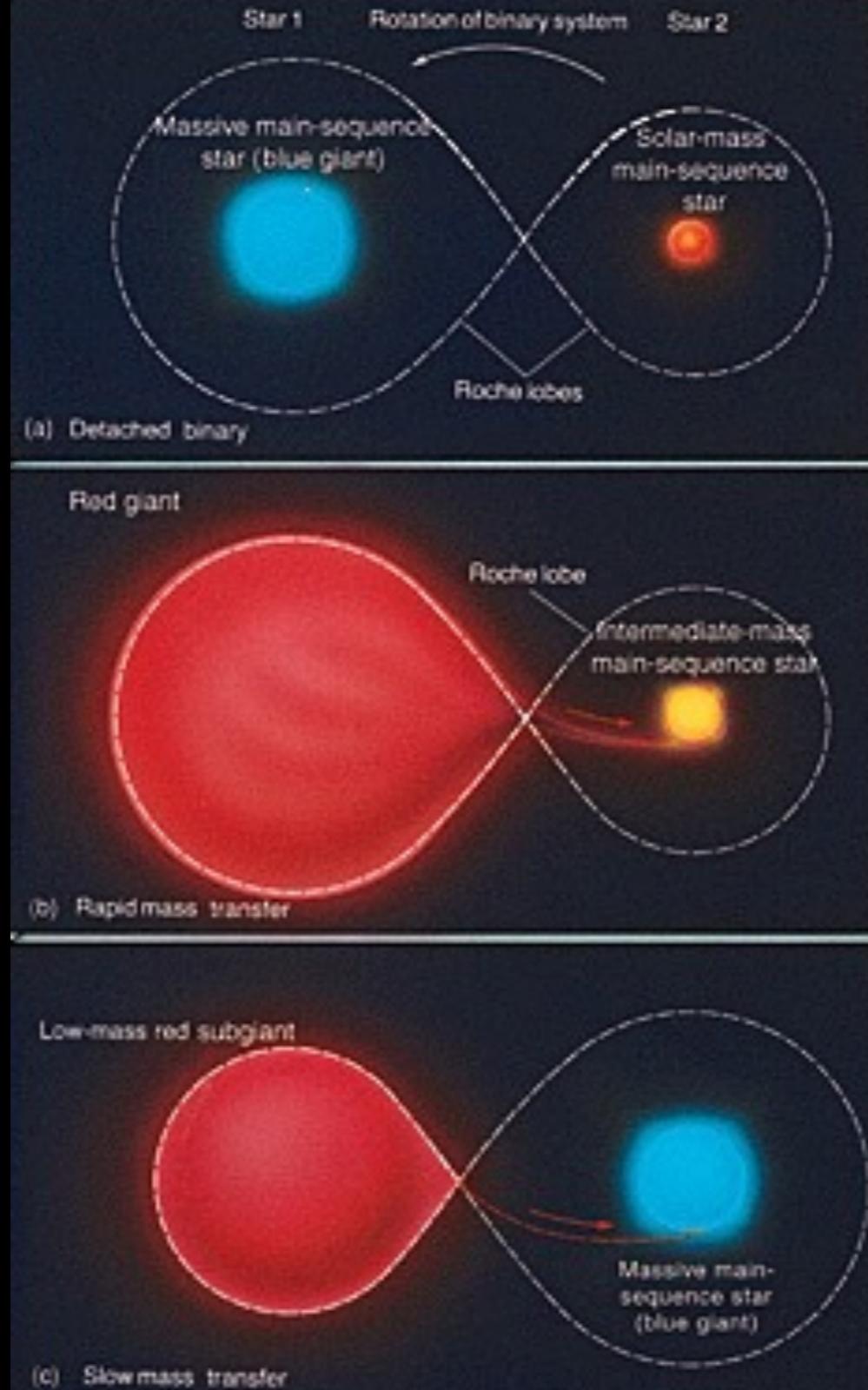
Matthew Bate
University of Exeter

Формирование звёзд



Изображение: «Astronomy Today 2E», Prentice Hall

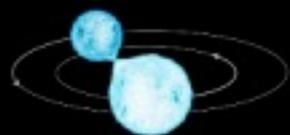
Эволюция Солнца



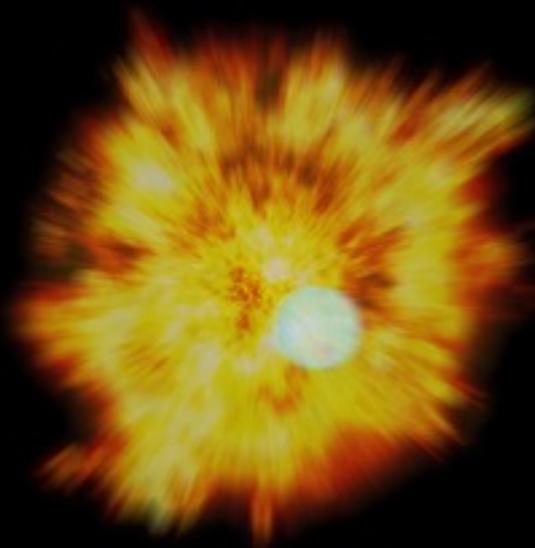
Изображение: «Astronomy Today 2E», Prentice Hall

Пример эволюции тесной двойной звезды

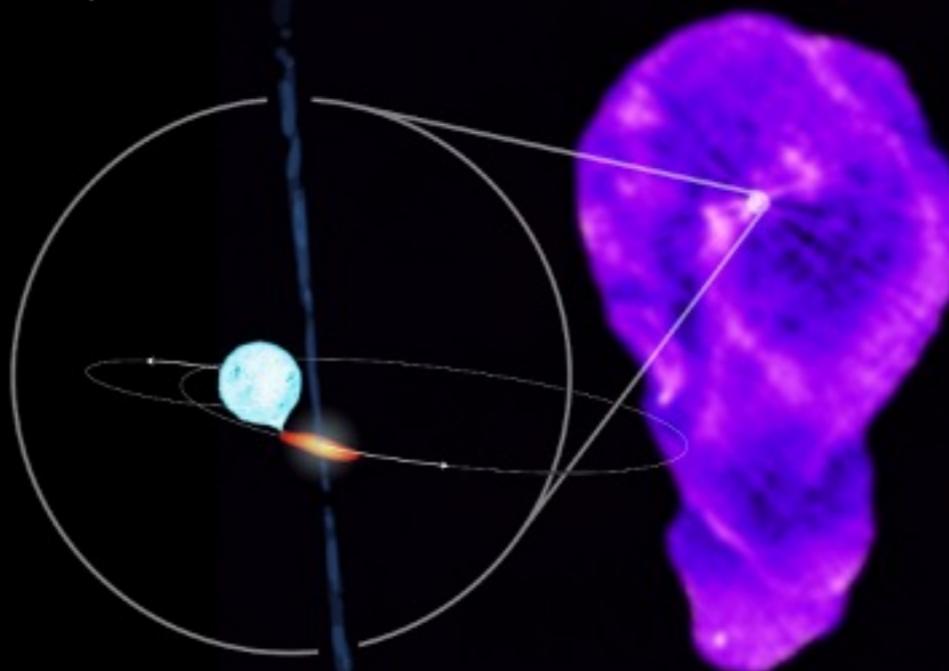
1) Two ordinary stars orbit each other for a few million years after they were born.



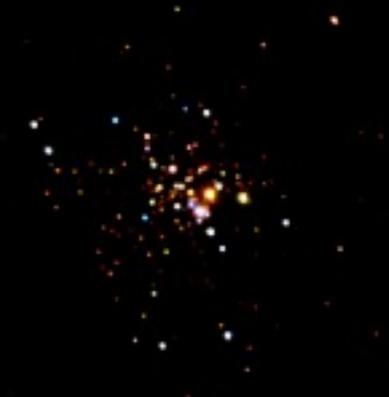
2) The initially more massive star runs out of nuclear fuel and explodes in a supernova, leaving behind a neutron star.



3) The neutron star and its companion in **Circinus X-1** orbit each other inside the short-lived residual glow of the supernova explosion. The explosion distorted the orbit to become elliptical. At closest approach, the neutron star is force-fed by its companion.



4) We observe **Circinus X-1** approximately 2,500 years after the explosion. Typical X-ray binaries are observed millions to billions of years after the supernova. Their supernova remnants have long since disappeared.



Old X-ray binaries in the globular cluster 47 Tuc observed by the Chandra X-ray observatory



more than 1,000,000 B.C.E.



Approx. 500 B.C.E.



2,013 C.E.



More than 1,000,000 C.E.

Time

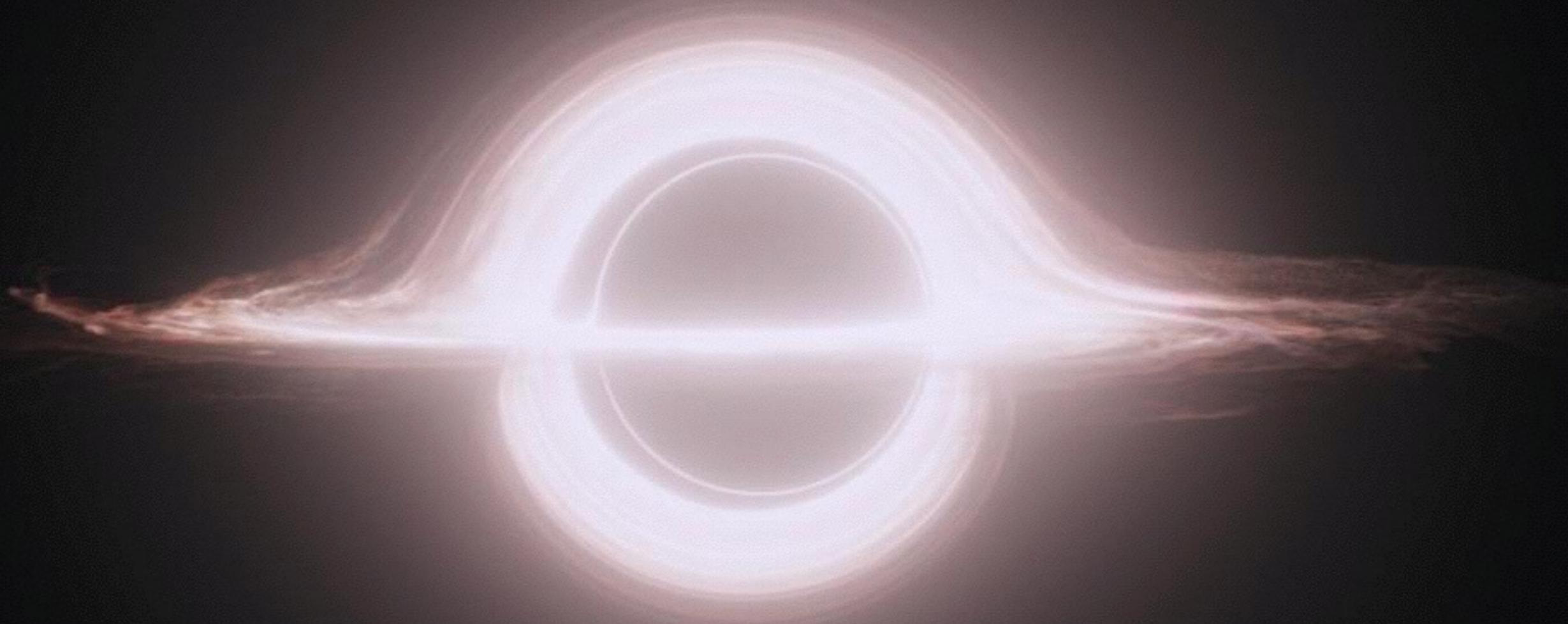
Изображение: Univ. of Wisconsin-Madison/S.Heinz et al

Эволюция Циркуля X-1



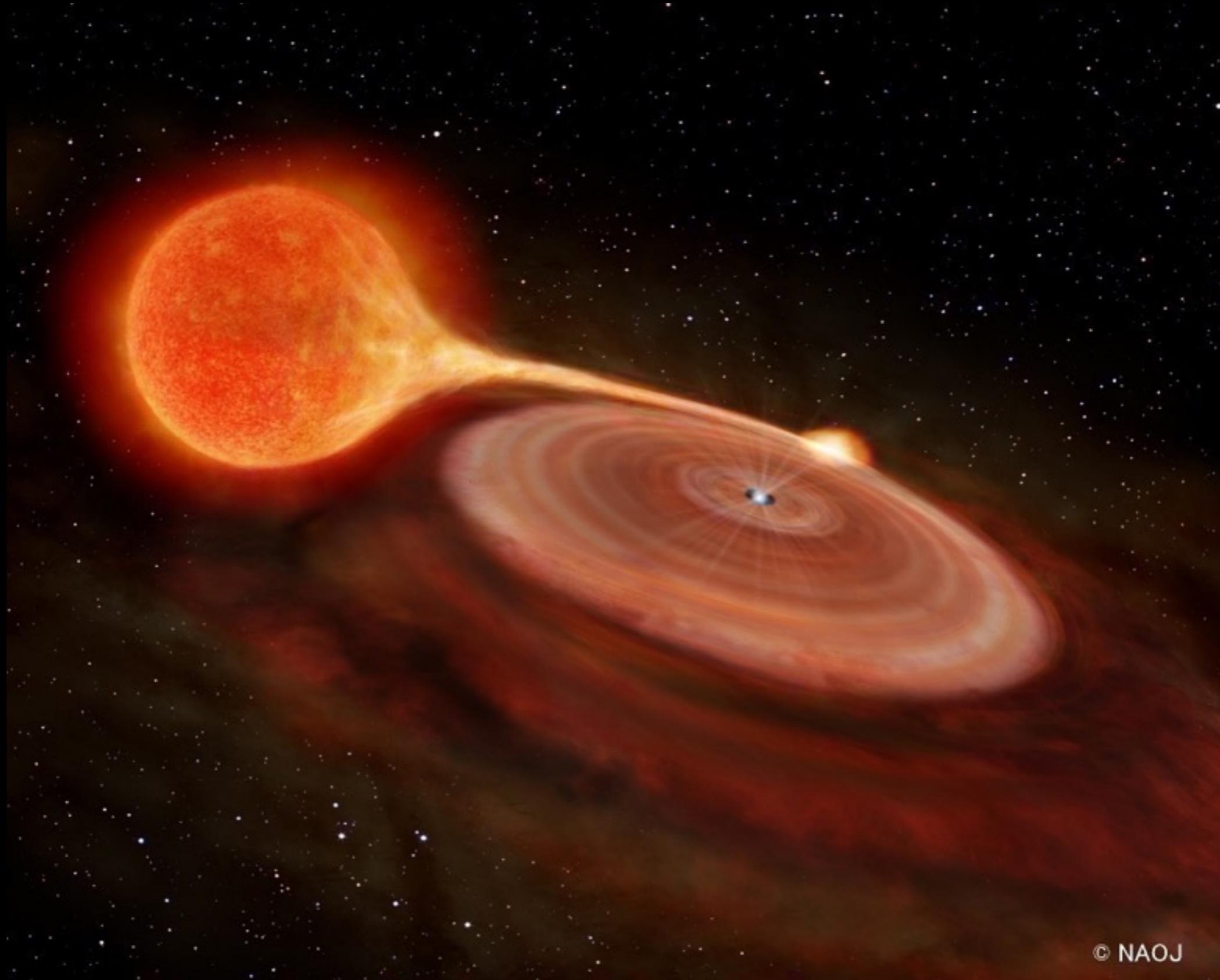
Изображение: NASA's Marshall Space Flight Center

Циркуль X-1



Изображение: Double Negative, Warner Bros.

Аккреционный диск



Карликовая новая



<https://www.flickr.com/photos/gsfcc/19115493618>

Видео: NASA Goddard Space Flight Center

Аккреция на чёрную дыру



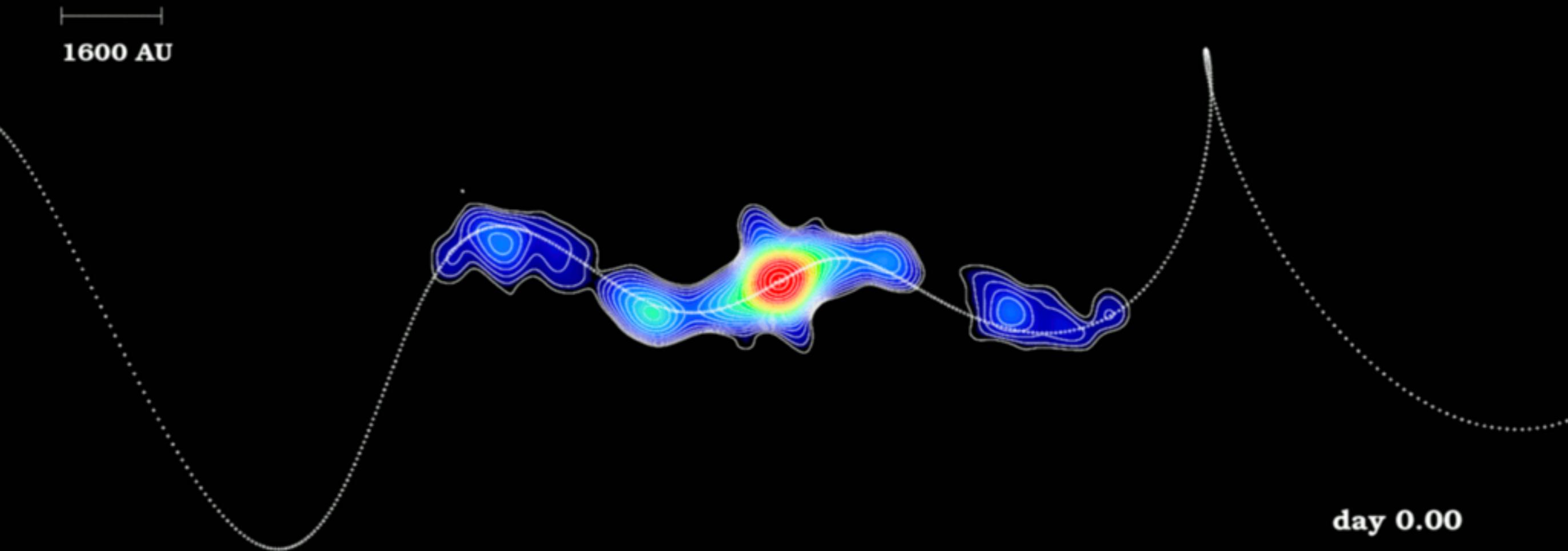
Изображение: NRAO/AUI/NSF, K. Golap, M. Goss; NASA's WISE

Остаток сверхновой



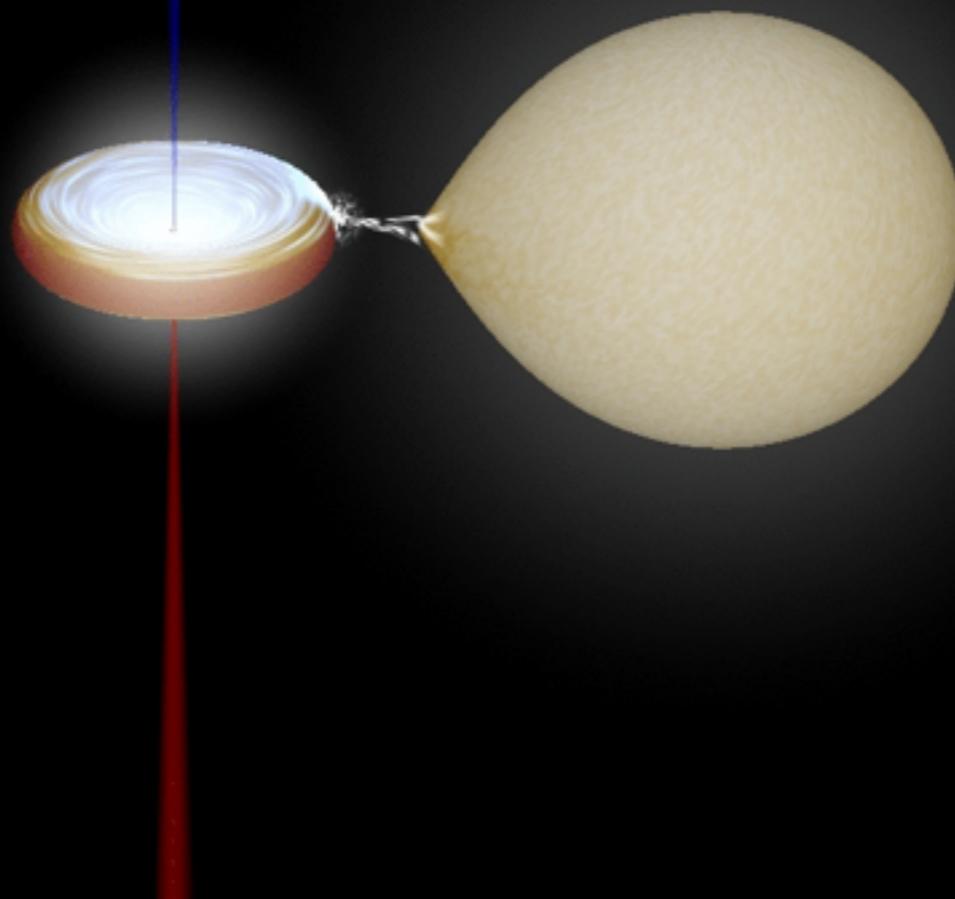
Фото: Т. Colson

Ламантин



Анимация: А. Mioduszewski et al., NRAO/AUI/NSF

Микроквazar SS 433



Анимация: Robert Hynes, Tania Burchell

Микроквazar SS 433

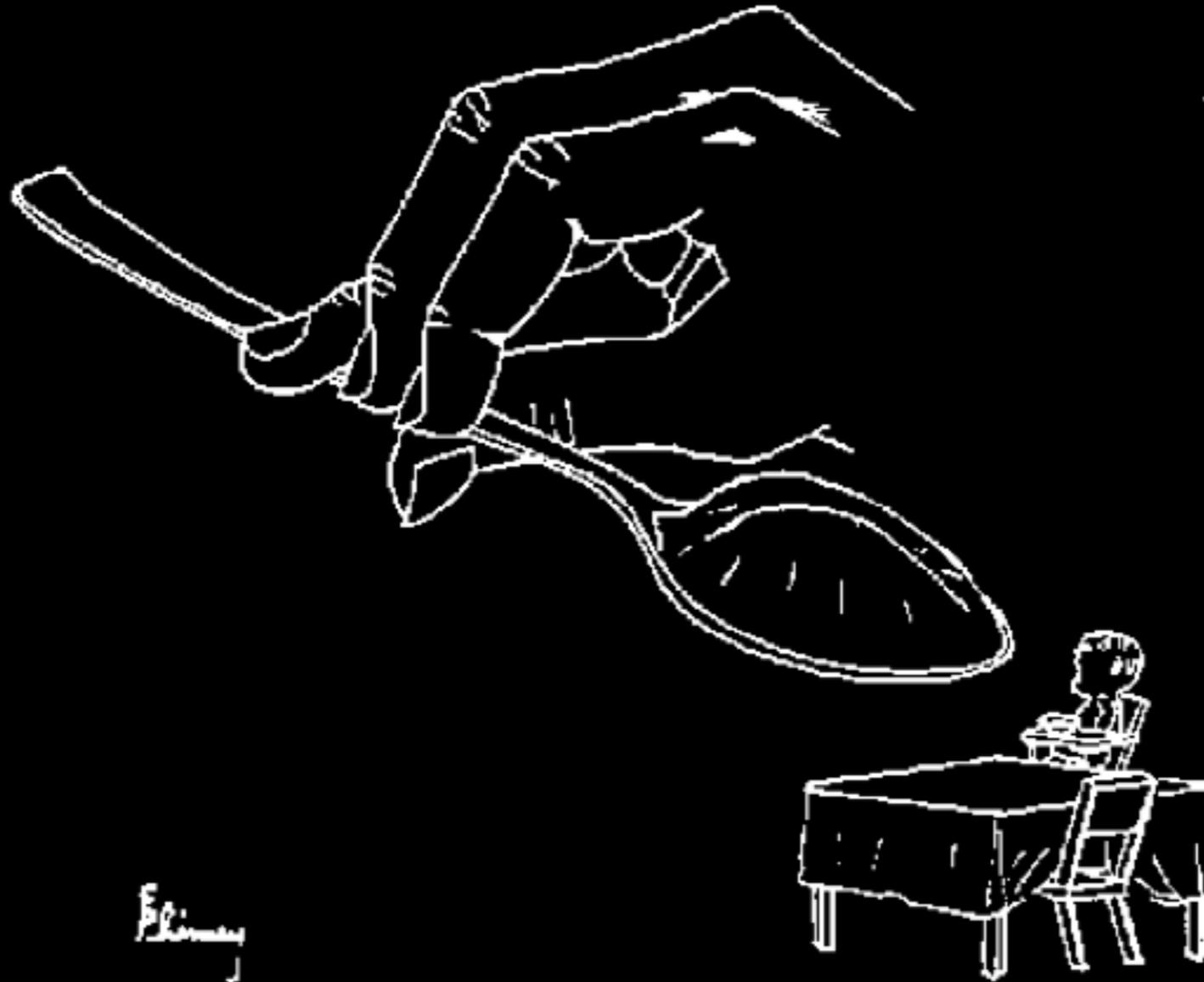


Fig. 1—The problem of feeding the monster: a large (angular momentum) spoon and a small (angular momentum) mouth. Hands and teeth (gravitational and magnetic forces, viscosity, . . .) are needed to guide and divide the food into morsels that can be metabolized during activity.

Сверхкритическая аккреция



<https://www.flickr.com/photos/gsfcc/4889085984>

Видео: NASA Goddard Space Flight Center

Новая — термоядерный взрыв
на белом карлике



<http://www.eso.org/public/videos/eso0943b/>

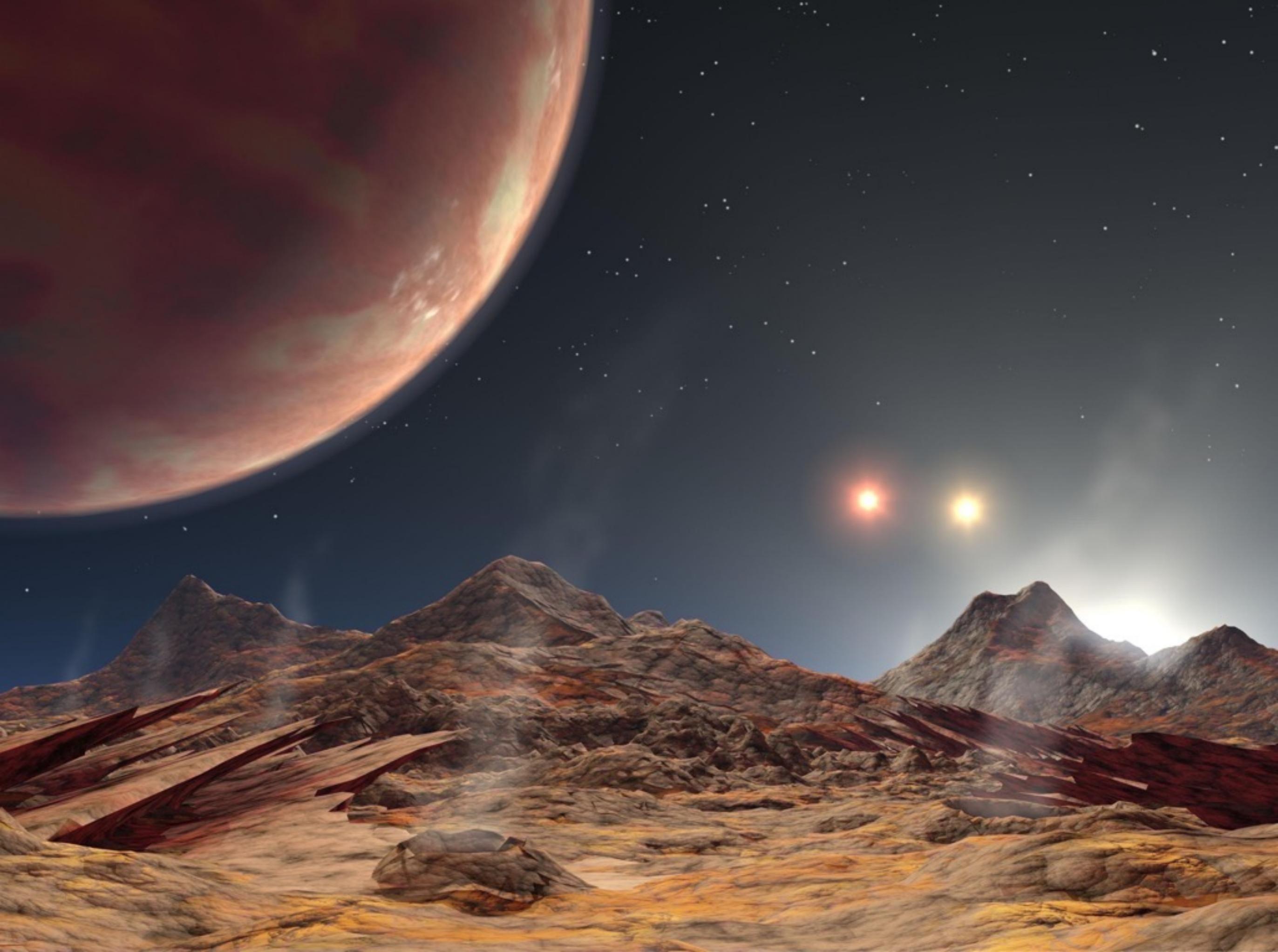
Видео:ESO/M. Kornmesser

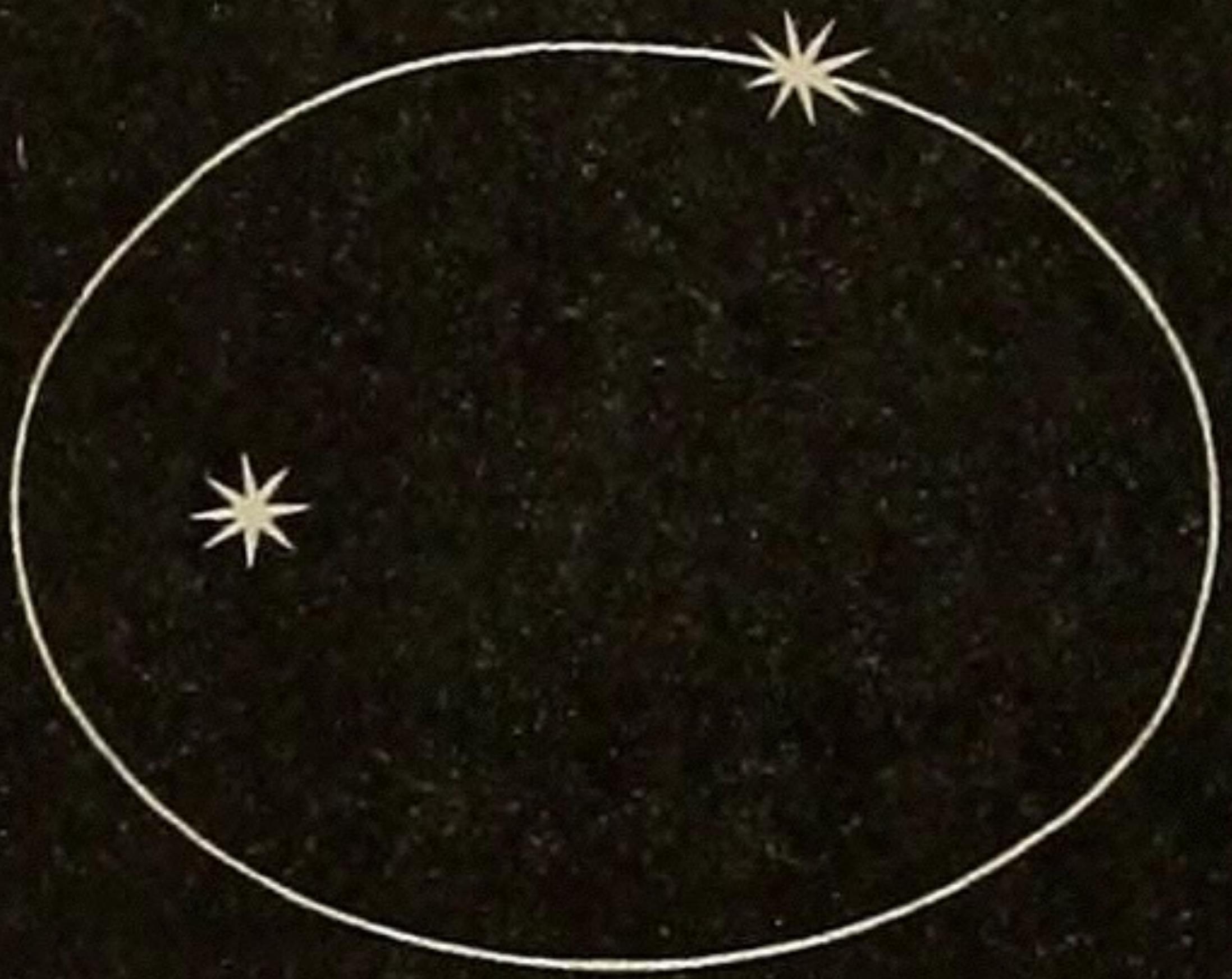
Сверхновая Ia — термоядерный взрыв
белого карлика



Изображение: NASA/CXC/Middlebury College/F.Winklerch

Остаток взрыва сверхновой Ia





Спасибо за внимание